

Comparison of RT codes via BBHRP

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BBHRP with different RT models

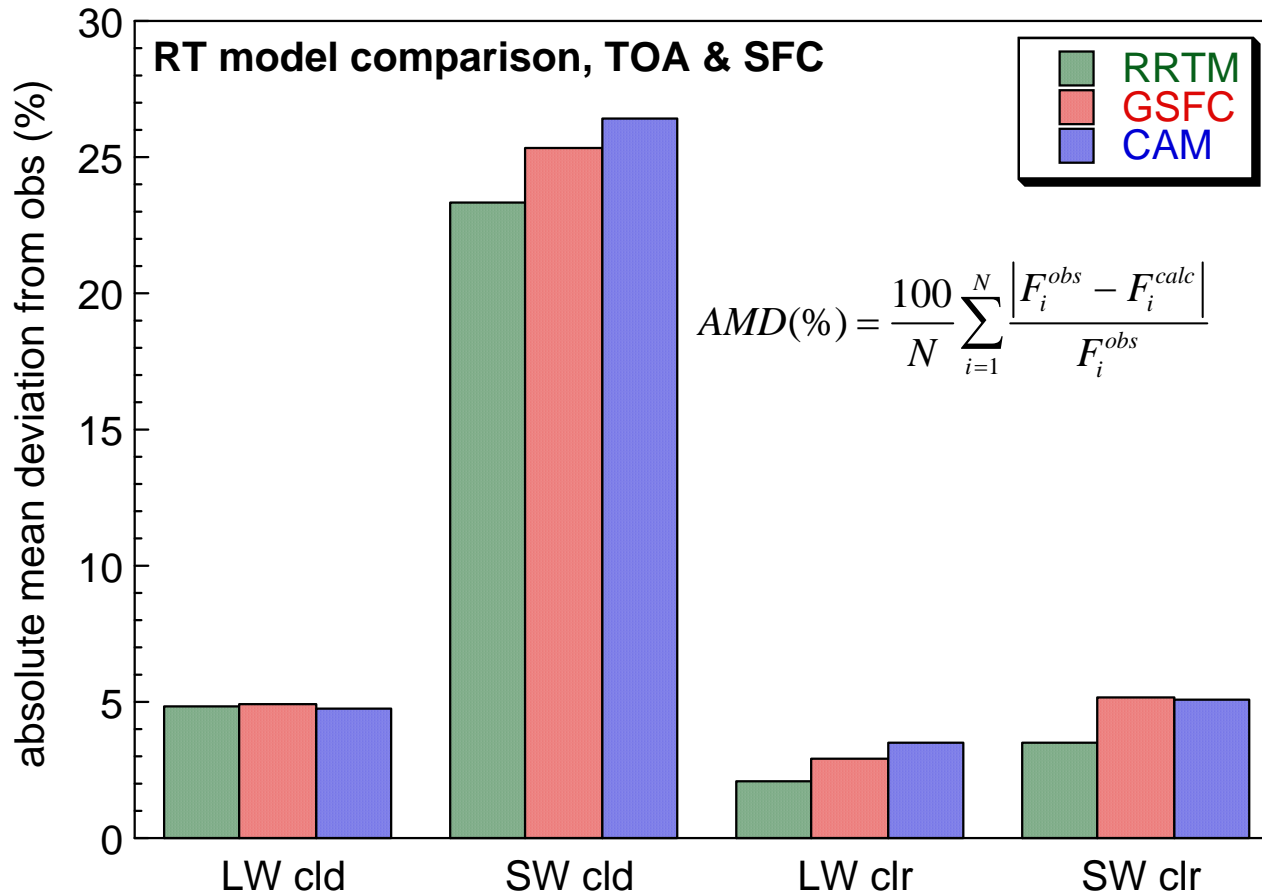
(see also Oreopoulos et al. poster)

Our goal: To learn when and why we succeed or fail to achieve radiative flux closure in BBHRP.

Our tools: The BBHRP dataset itself and radiative fluxes calculated as in BBHRP (aka our “shadow” dataset) with two pairs (SW and LW) of additional RT algorithms: from CAM3 and from GSFC’s fvGCM. BBHRP uses AER’s SW and LW RRTM codes.

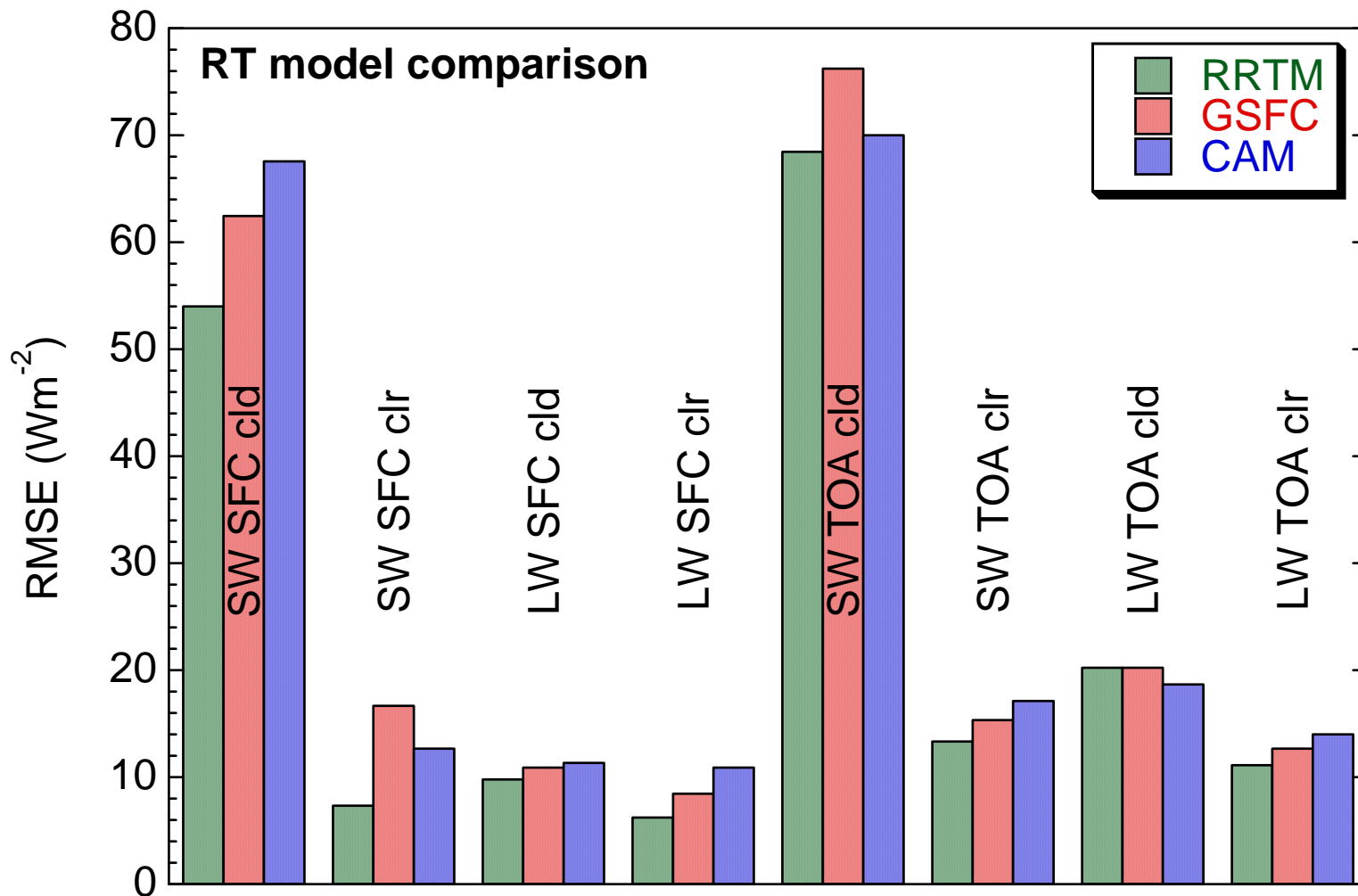
How we learn: If the RT models tend to agree, but disagree with the observations for particular types of conditions the input is suspect; if on the other hand, for the same conditions only some models fail, these models are suspect.

BBHRP with different RT models



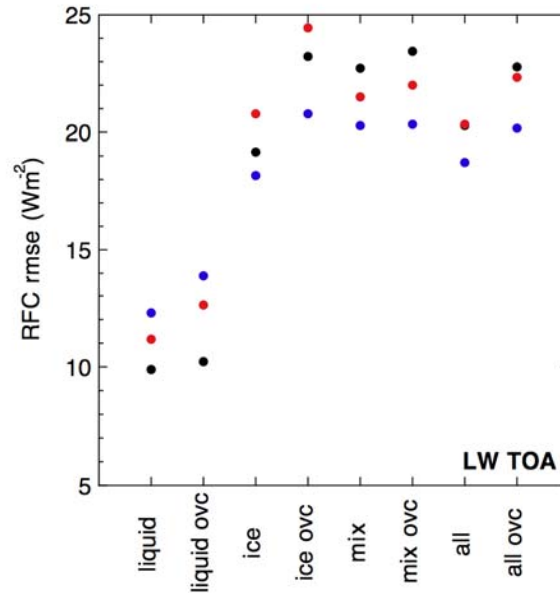
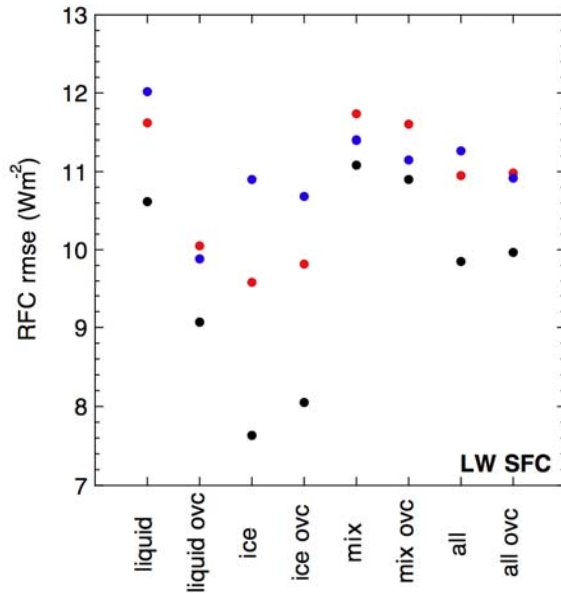
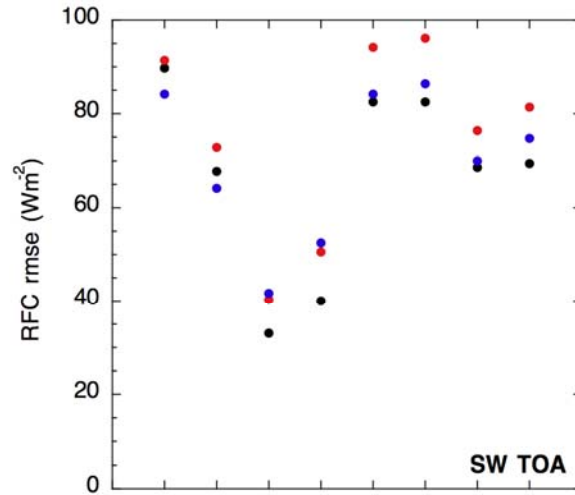
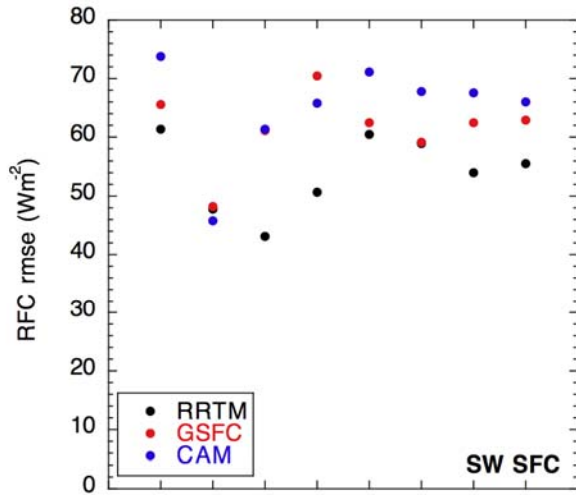
BBHRP with different RT models

(more details: SFC vs. TOA)



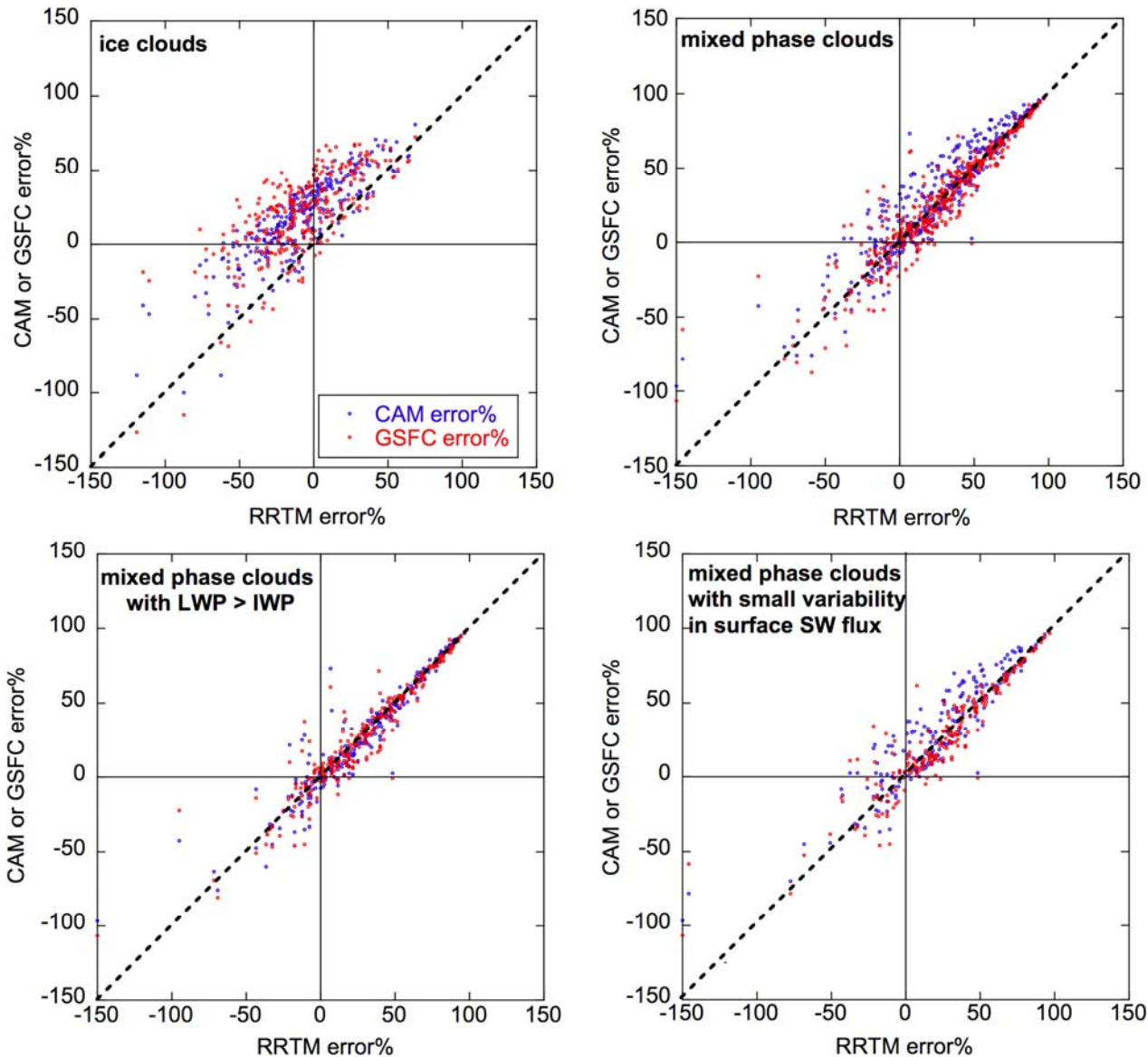
BBHRP with different RT models

(more details: cloud type)



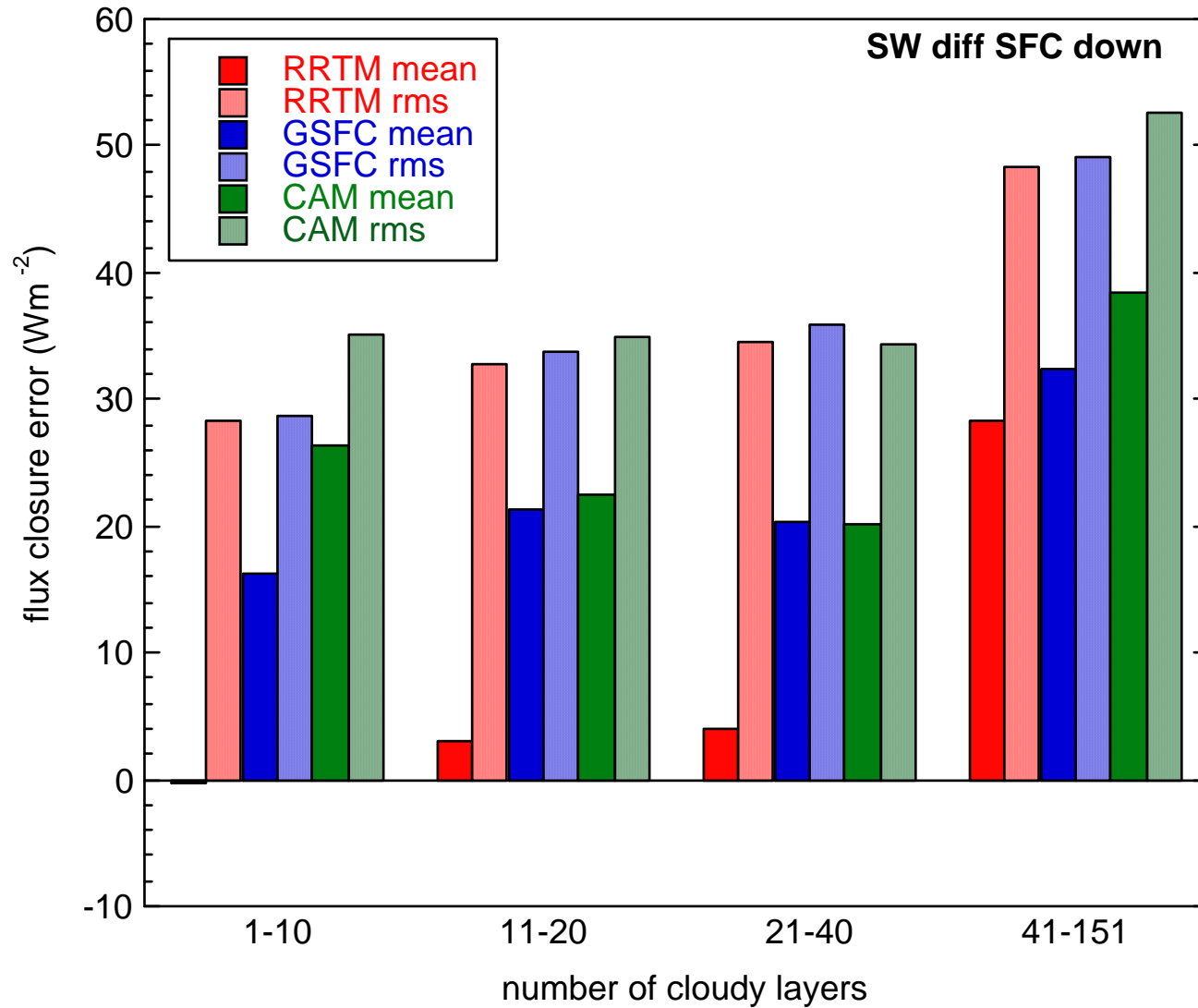
BBHRP with different RT models

(more details: ice and mixed clouds)



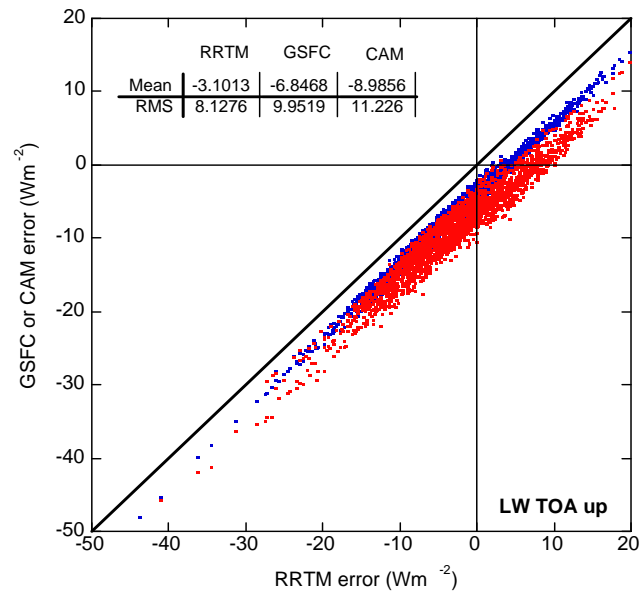
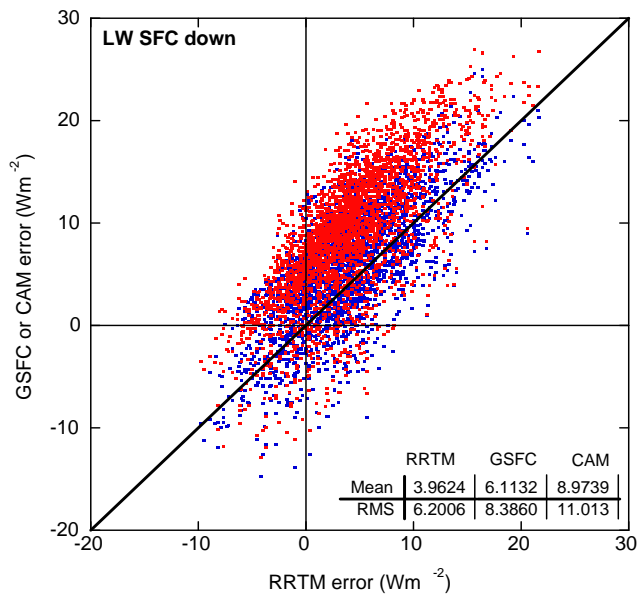
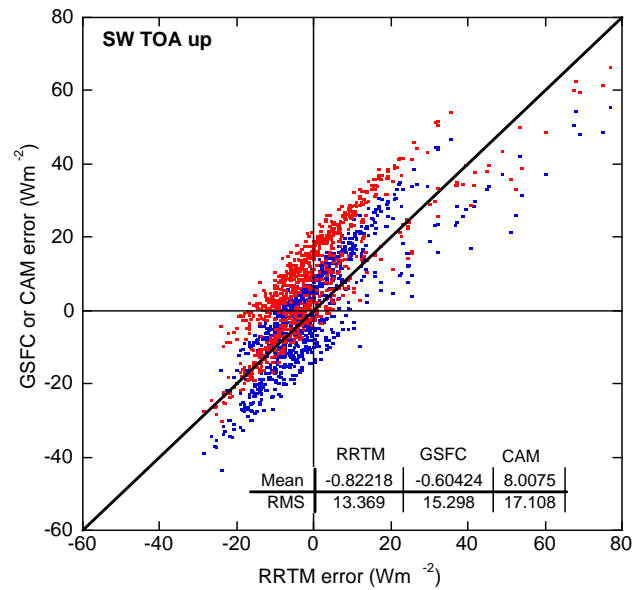
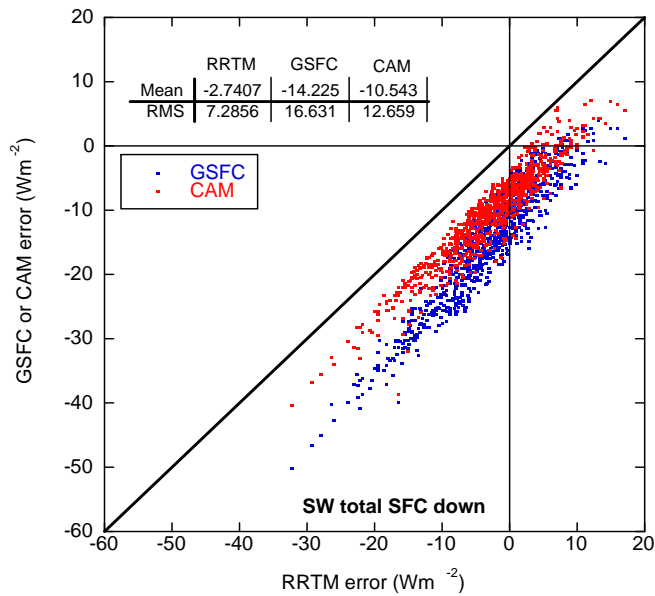
BBHRP with different RT models

(cloud complexity)



BBHRP with different RT models

(more details: clear sky scatter)



Main findings

(while a lot of questions remain)

- There is broad consistency between the BBHRP (RRTM) closure errors and those of the “shadow” dataset (GCM codes), pointing to problems in the values or interpretations of the input.
- Inter-model inconsistencies are greater for ice and mixed-phase clouds.
- Many large SW SFC closure errors are associated with very thick clouds
- For clear skies the GCM codes always produce higher values than RRTM for SW down at SFC and OLR (needs to be investigated).
- RRTM performs overall better than the GCM codes.
- When all radiation budget components are accounted for (SW & LW, TOA & SFC), under all conditions, the overall flux closure error is ~10%. This is driven largely by the LW, but is still remarkable.

Additional slides

