

Column Water Vapor From Diffuse Irradiance

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Abstract: A possibility of measuring water vapor column from diffuse irradiance, and thus the extension of optical retrievals to cloudy days, was investigated. The data from the rotating shadowland systemorimoters (RSS) during in whater deployment at the North Spee of Alakas (NSA) site is used. The initial analysis covers 20 days in March 1999 that include clear, partly cloudy and overcast days. During these days water vapor column according to the NSA site's microwave radiomater (MWR) varied between one and five mm. The diffuse irradiances in the 820-am and 940-am water vapor absorption bands are compared with water vapor column obtained from the MWR. While these irradiances do not correlate well with the water vapor column, we found that by using the diffuse transmittance in the 760-am oxygen absorption hand to obtain an effective air mass we could devive a method that greatly increases the correlation. The end result is a correlation of 0.07 and 0.95 between MWR and diffuse RSS retrievals using \$200-am and 940-am transmittances respectively. This concursion greatly is shoot on empirical data correlation. The end result is a correlation of U.97 and U.95 between MWK and armise RS5 resurvas usuage \$20-m and 940-m matmaintitunes, expectively. This encouraging result is based on empirical data analysis only. It implies that the diffuse irradiance may contain sufficient information to retrieve water vapor column. Bowever, we must analyze broader data sets to ensure that the resulting high correlation in these data is not predicated on the limited climatological conditions experienced in March 1999 at the NSA ""

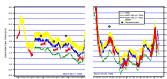


Figure 1 Microwave and RSS direct irradiance retrievals

Introduction: The intensive observation period (IOP) was conducted during the late winterlearly spring of 1999 at the North Slope of Alaska (NSA) ARM site near Barrow, Alaska, to compare various techniques to retrieve water vapor column in dy conditions. The rotating shadowals apectronalometer (RSS), deployed by the Atmospheric Sciences Research Center, was the only sun photometer taking part in this resperiment. In addition to the 23 soft EM 2MW, NOAA Environmental Technology Ladoratory and NASA Goddard Space Flight Center deployed two radiometers covering the 183 GHz band. The results of retrievals with microwave endioneters can be found in Racette et al. [2000] and Westwater et al. [2000] and Westwater et al. [2000] and Westwater et al. [2000] and offered them reasonable of the comparison between the optical and microwave methods was limited to 11 days of clear and semi-clear weather when solar direct beam measurements and retrievals could be performed. Fig. 1, reperioded from Kiedron et al. [in print], shows a good agreement between optical and microwave measurements.

The RSS provides continuous direct traditions expect on over 512 pixels covering the wavelengths

Kiedron et al. [in print], shows a good agreement between optical and microwave measurements.

The RSS provides continuous direct irradines spectra over 512 pixels covering the wavelengths
350–1080 nm (see Harrison at al. 1999). This allows us to derive better estimates of transmittance at
neighboring wavelengths. Then a good estimate of irradiance at he absorption band, as if there were no
auter, is calculated. After the baseline is removed from the irradiance, the transmittance at the water
absorption band is greatly desensitized to aerosols and radiometric calibration uncertainties. This method
allows one to retrieve water vapor column even when the sun is partfally obscured, or in the presence of
fog, when only a fraction of direct beam can reach the instrument. But beyond some degree of direct beam
cutation, the low signal-to-noise ratio makes retrievable impossible.

The addition to the direct horizontal irradiance, the RSS also measures the diffuse irradiance. The
latter is a distinct on the contraction of the contra

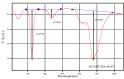




Figure 2 Diffuse spectrum

Methodology: We try to find a function of diffuse irradiance that predicts the water vapor column retrieved by the MWR. More precisely, we concentrate on only two spectral elements of the diffuse irradiance; one at the water absorption band near 940-nn or 820-nn and the second at the oxygen absorption band near 760-nn. Muthematically speaking, we try to show that two spectral elements from diffuse irradiance produce a resultant that correlates with water vapor column from the MWR measurement. In other words, we create an estimator F of water vapor column w:

$$w = F(I_{dif}(\lambda), I_{dif}(760), \overline{x}) \tag{1}$$

where λ =940-nm or λ =820-nm and \overline{X} is the vector of parameters that are derived through the process that

where $\lambda \sim 940$ -nm or $\lambda \sim 820$ -m and λ is us vector to promote minimizes residuals $\tau = 0$ -magnet. To eliminate sensitivity to aerosols and radiometric calibration errors, transmittances, instead of irradiances, at water and oxygen absorption bands are used. The irradiance is divided by the extraterrestrial and the baseline is removed. In Figure 2 we show diffuse irradiance divided by the extraterrestrial irradiance. The anchor points for each hard 8 seeline are indicated. The value of the baseline at absorption wavelength λ is interpolated using values at anchor points λ_s and λ_c with the formula

$$B(\lambda) = I_{dif}(\lambda_1) \left(\frac{I_{dif}(\lambda_2)}{I_{dif}(\lambda_1)} \right)^{\frac{\lambda_2 - \lambda_1}{\lambda_2 - \lambda_1}} \eqno(2)$$

Then the transmittance at absorption wavelength is given by

$$T_{dif}(\lambda) = I_{dif}(\lambda)/B(\lambda)$$
 (3)

The inclusion of irradiance at the oxygen absorption band has the following rationale. The oxygen profile and its total column are invariant. Thus the absorption at oxygen band is dependent on geometry only and independent of variable atmospheric constituents and parameters. By incorporating the diffuse irradiance at 760nm we hope to compensate for the effects of varying pathlength with different atmospheric conditions as well as with the varying sun position. For the same reason high resolution measurements at 760nm band have been used and proposed to estimate the optical photon pathlength (Veifed et al. 1998.) Pellishicker at al. 1998.] Aim and tharrion [1999]. The estimator I will have no explicit depredaction of the contraction of the contr

$$m = g(T_{dir}(760))$$
 (4)

Once the function g() is established, we define the effective diffuse air mass as

$$m^* = g(T_{dif}(760))$$
 (5)

Next we try to find the functional relationship between the diffuse transmittance and the product of the w vapor column and the effective diffuse air mass

$$w_{MWR} \cdot m^* = h(T_{dif}(\lambda))$$
 (6)

$$= \frac{h(T_{dif}(\lambda))}{a(T - (240))}$$
(7)

Curve of growth model: We used Moskalenko's [1969] model of the curve of growth for both functions g/s) and s/s, however, we must emphasize that the selection of the model function is not essential to our task. In fact, to establish correlation between diffuse transmittance and water vapor column, functions g/s) and s/s p/s contained to establish correlation between diffuse transmittance and water vapor column, functions had not some possible of the source of

$$\ln T = -k \cdot x^{f(x)}$$

where f(x) varies between I and I/2. We chose f(x) in the following form

$$f(x) = 1 - a \frac{x}{x + b}$$

with two positive parameters a and b. Always g(0)=1 and $g(\infty)=1$ -a. However we did not limit a to assure the physical condition $g(\infty)=0.5$. To obtain values of x=g(T) or x=h(T) eqn. (8) was solved numerically with

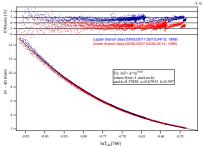


Figure 3 Curve of growth at 760nm

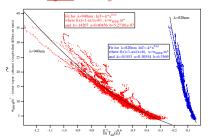


Figure 4 Diffuse curves of growth

Results: The 20-day data set totals over 12.000 diffuse and direct irradiance spectra. Among the data six Results: The 20-day data set totals over 12,000 diffuse and direct imidiance spectra. Among the data six days were clear. We added data from two clear days in April to generate the curve of growth g/f from the direct transmittance at 760-mn. In Figure 3 the air mass excording to Rasten [1965] is plotted against the oligorithm of transmittance. The residuals show several systematic tendencies of which one remains unexplained. We do not know why data from four days cluster above the best-fit curve and from the remaining four days below the curve. The sudden step change in residuals for buy air masses is due to non-linearity in the RSS's NMOS array after exposure changes. The increase of residuals for large air masses is due to the lower signal-no-noiser neith of large solar zerulta neighes. The morning and afternoon branches for each day coincide implying that RSS's time keeping is accurate. In summany, the residuals are still within transmittances at R300 am and 90-0m are devired in Figure 4. It is evident that the outliers are not random. This suggests a need for an additional parameterization. Nevertheless, for each data point we calculate two estimates of water veryon column with equ. (7). The results are presented in two correlation plots in Fig. 5 and in Fig. 6 as a function time.

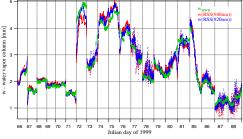


Figure 6 MWR and diffuse retrievals

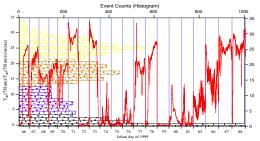


Figure 7 Direct-to-diffuse ratio

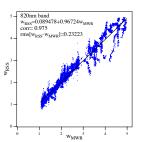
Discussion: The correlation of 0.97 and 0.95 between MWR and diffuse RSS retrievals using 820-nm and 900-nm transmittances, respectively, can be considered very high. We speculate that the retrievals with 820-nm are slightly better because of the fact that the 820-nm photons travel a more similar path to the 760-nm photons than the 340-nm photons. The retrieval errors have a random component with zeon mean

820-ms are slightly better because of the fact that the 820-ms photons travel a more similar path to the 70-ms photons have a human made with the 820-ms photons travel a more similar path to the 70-ms photons have a readon component will zoro mean and a systematic component.

The medium error is particularly large on clear days jul-77, 78, 75 and 85 (see Fig. 6). This is because the 70-ms particularly large on clear days jul-77, 87, 87 and 85 (see Fig. 6). This is because the 70-ms particularly large one clear days jul-77, 78, 87 and 85 (see Fig. 6). This is because the 70-ms particularly large one clear days jul-77, 87, 87 and 85 (see Fig. 6). This is because the 70-ms particularly large one clear days days. This error can be reduced by averaging data points or by implementing a change in RSS firmware to perform the blocked measurement that must be kept below the detector statution point. In this paper, we are chefully concerned with the systematic errors as their magnitude determines the validities of the particular control of the particula

$$r = \frac{1}{\cos(SZA)} \frac{T_{dir}(781)}{T_{rr}(781)}$$
(10)

 $\cos(3.62) \cdot 3_{\rm H}(183)$ at a non absorbing wavelength of 781-mm to distinguish between degree of overcast. at a non absorbing wavelength of 781-m to distinguish between degree of overcast. The histogram suggests a division for Figure 7, along with the ratio r_i a histogram of r_i values is plotted. The histogram suggests a division for Figure 7, along with the ratio r_i and r_i was a second proper of the content of the value of r_i was a second proper of the content of r_i . While the improvement occurs, it is not large because the majority of all the outliers are within the same group $(0x-r_i)$. These are the points from the overcast days and points from the clear days at very large solar canding adjust the formation on the solar parall maglie into the model to distinguish between the two types of low r_i points. We can imagine further improvements by adding memory into the model. The presented model does retrievals institution/oxid parallel values of the content o



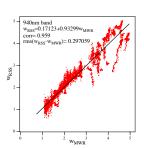


Figure 5 Correlation plots

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